

EAVN and SKA

A large radio telescope dish is shown at night, illuminated from below. The dish is white and mounted on a complex metal structure. The background is a dark blue night sky filled with stars, with the Milky Way galaxy visible as a bright, hazy band of light stretching across the upper left portion of the frame. The foreground shows the dark silhouettes of trees and a hillside.

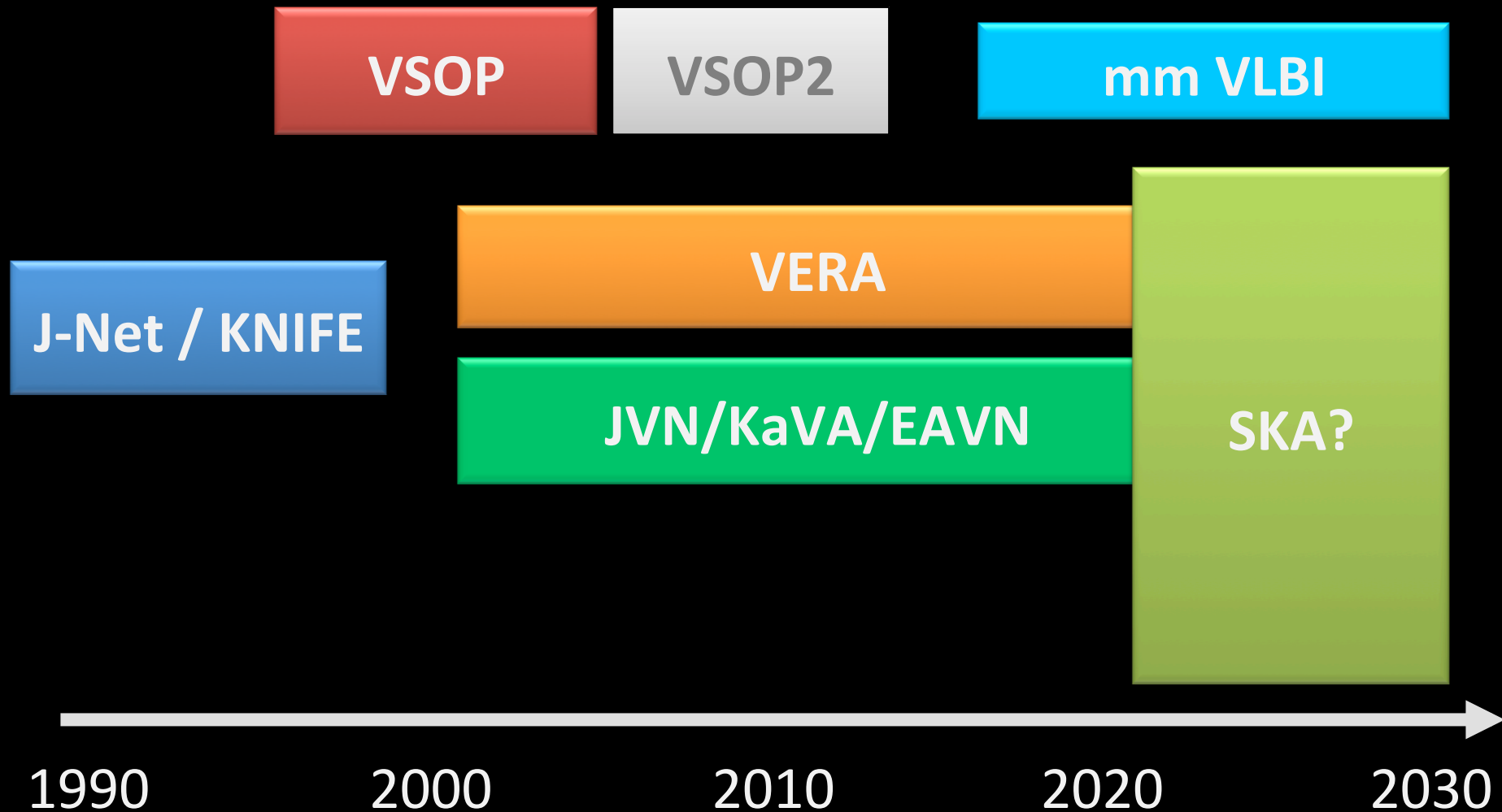
Mareki Honma

Director, Mizusawa VLBI Observatory, NAOJ

水沢VLBI観測所の目標(FY2015)

1. VERAによる観測を継続して研究成果を挙げる。
2. 大学間連携・国内VLBI観測網の中核局として大学を支援し、大学間連携VLBIで研究成果を挙げる。大学連携VLBI観測の体制および事業内容を検証する。
3. VERAの観測が7年後に終了することを想定し、日本の電波コミュニティ等の意向に沿う形の水沢VLBI観測所の将来計画を検討する。
4. 東アジアVLBI観測網を運用し、研究成果を挙げる。
5. 中央標準時の決定と現示を行う。

Brief history of VLBI at NAOJ



Announcement on mm VLBI

- VLBI with ALMA will be open from **Cycle-4** (risk-share mode)
- **Band 3 (7mm) and Band 6 (1.3mm)** will be available
- **Two proposals needed** (to the array and to ALMA)
- The array proposal for Band 3 should go to **GMVA**. The proposal for Band 6 should go to **NRAO/EHT(TBD)**.
- The array proposal deadline will be **1 Feb 2016** (at least for GMVA)

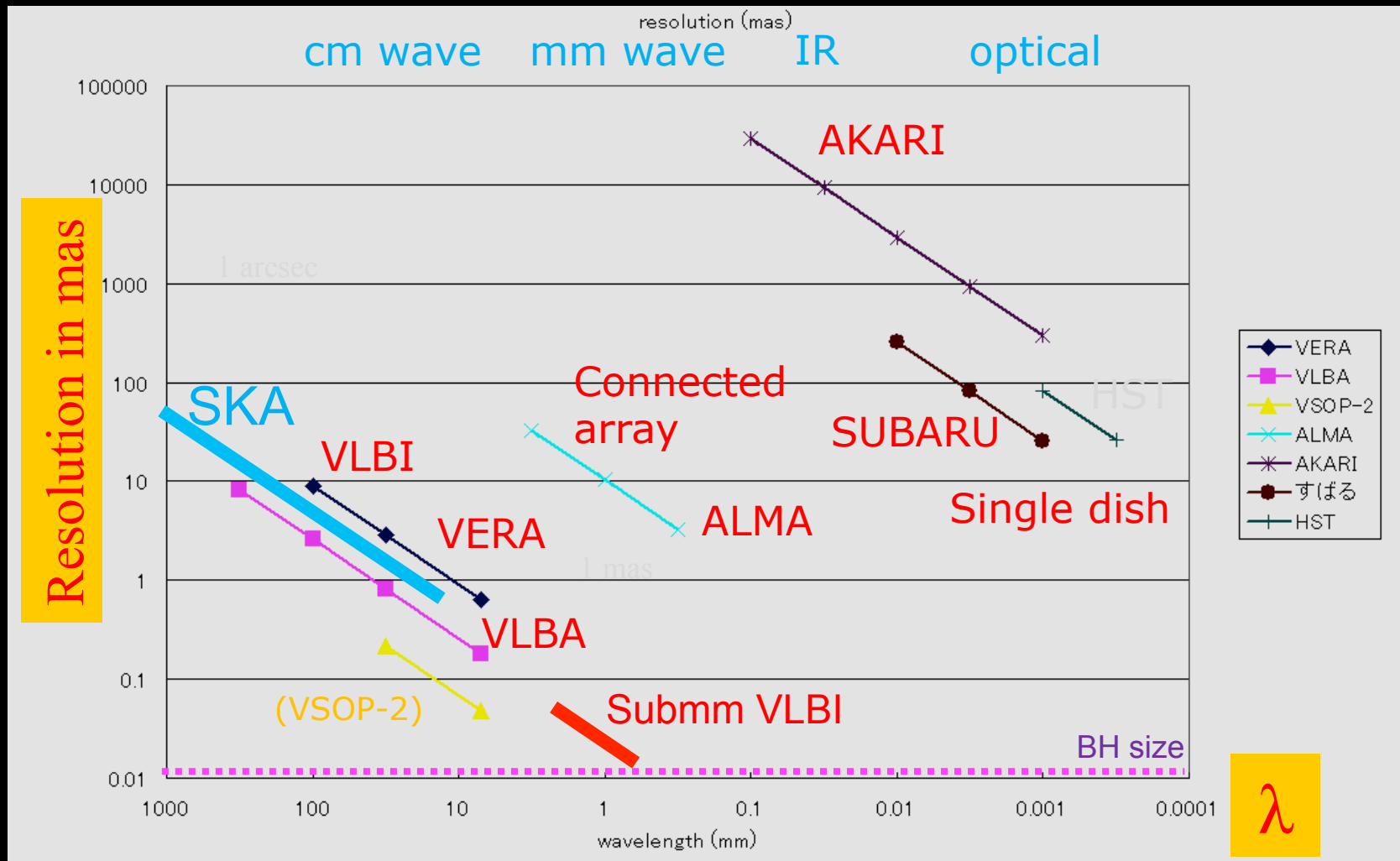
Radio astronomy in 2020's and beyond: SKA era

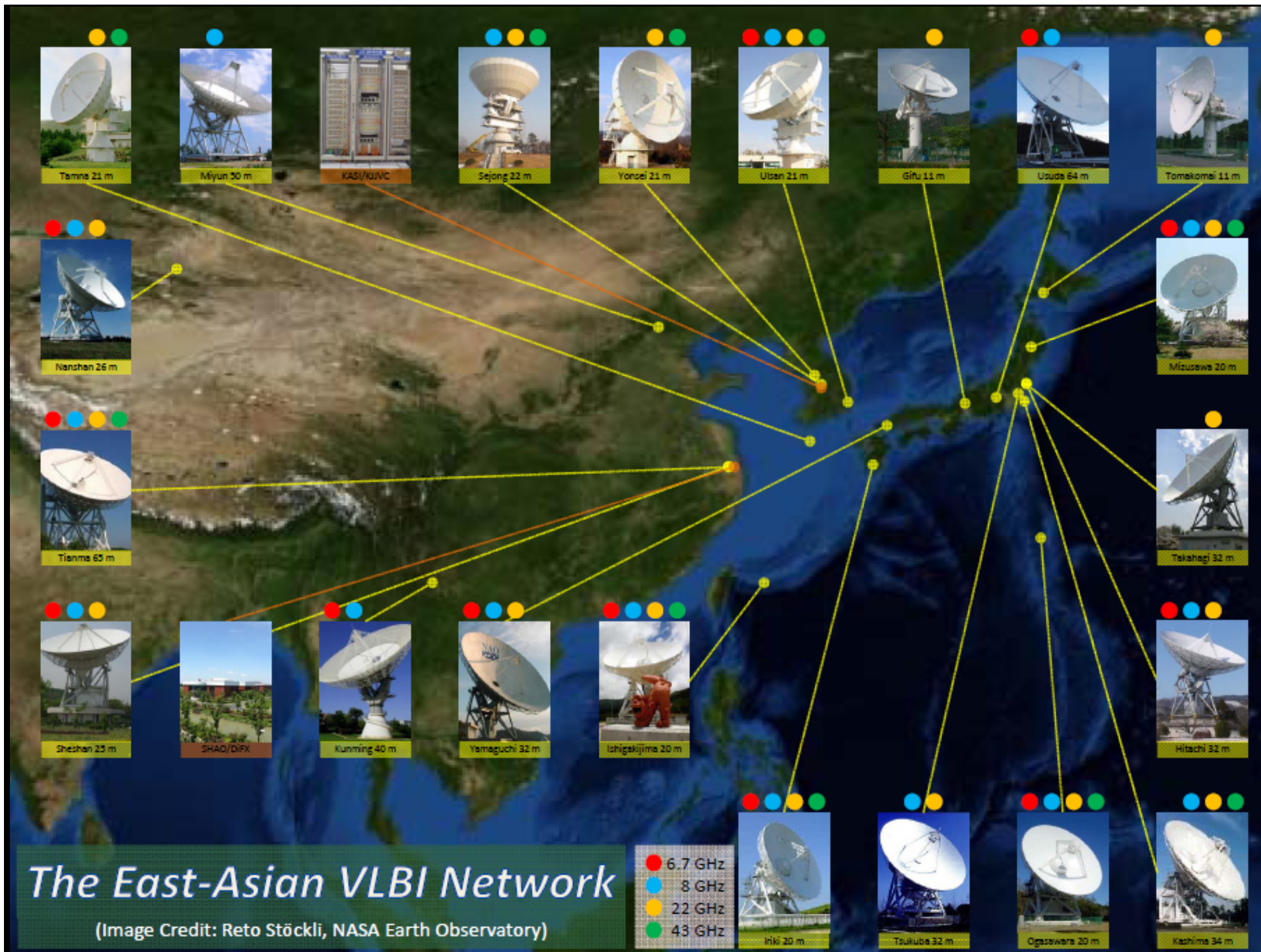
- SKA (**Square Kilometer Array**)
 - Next-generation big-project in radio astronomy
- Possible synergies with VLBI
 - technical overlap (interferometer)
 - frequency overlap, science overlap



Wavelength - Angular-resolution

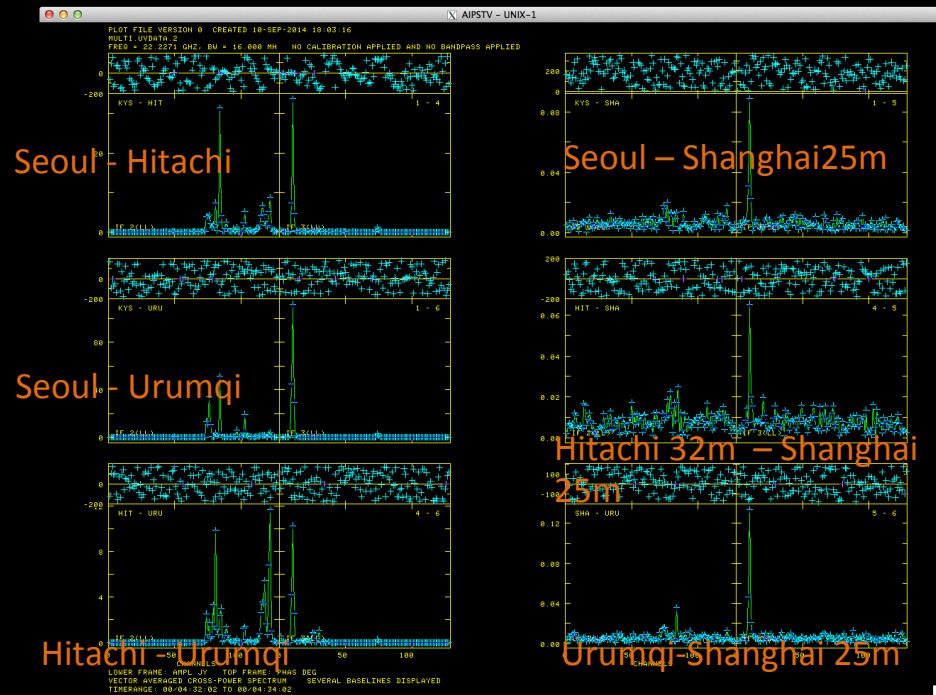
- angular resolution: $\theta \sim \lambda / D$





EAVN VLBI fringe tests

- 2014 January (3rd fringe test)
- Source: W49N (22 GHz maser)



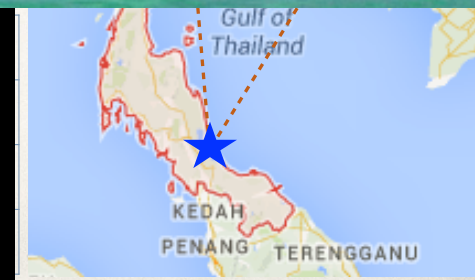
EAVN "Tiger Team" (Test observation team)
face-to-face meeting at Shanghai in July 2014

A plan for new telescope in Thailand

- NARIT has a plan to build TVN.



**KaVA Science WG in Krabi
21-24 Feb 2016**



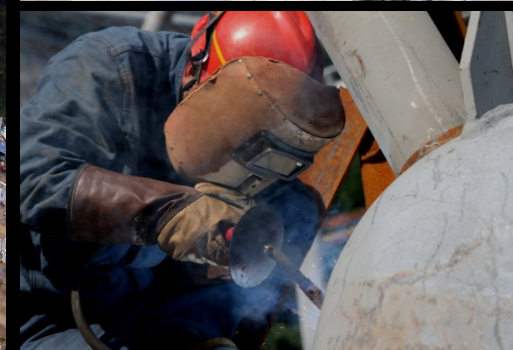
Comparison of SKA, EAVN and other array

	EAVN	EAVN +FAST	EAVN +QT,FAS T	EVN	VLBA+ GBT+ VLA	Global VLBI (EA/Eu/ US)	SKA-1 (mid)	SKA-2 (mid)
Operating from	2018?	2022?	2022?	Operating	Operating	?	2023?	2028??
Max. Baseline (km)	5000	5000	5000	2000-10000	8000	10000	150	3000
Collecting area (m ²)	15000	86000	96000	20000	26000	61000 ~ 142000	32600	440000

FAST 500m corresponds to an effective diameter of 300m

QT: Planned new 110m in Urumqi

建設中のFAST, 2016年秋完成予定



Science cases

- AGNs
 - Masers, SFR, AGB stars
 - Galactic structure
- +
- Pulsars
 - Radio transients
 - Cosmology
 - SETI (!?)

Proposal

Shouldn't we seriously consider what to do with VLBI in SKA era ?