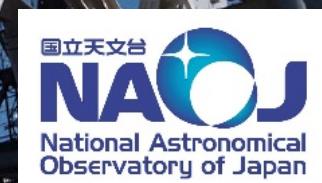


VLBI位置天文観測に準拠した、 密度波理論の検証

**Test of the Density-wave theory
Using VLBI astrometry results**

Sakai et al. (2015)

VERA

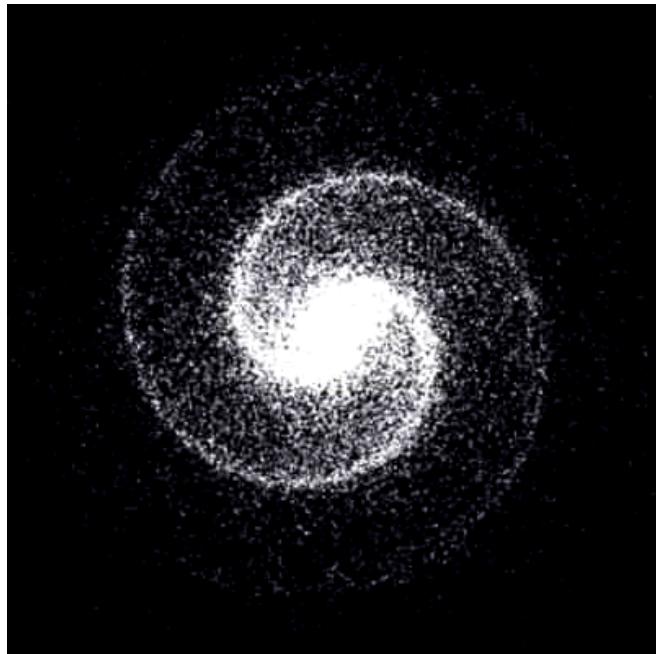


Nobuyuki Sakai, National Astronomical Observatory of Japan
Dec. 24, 2015@Toyo Univ., Tokyo

Theories for the Spiral arm

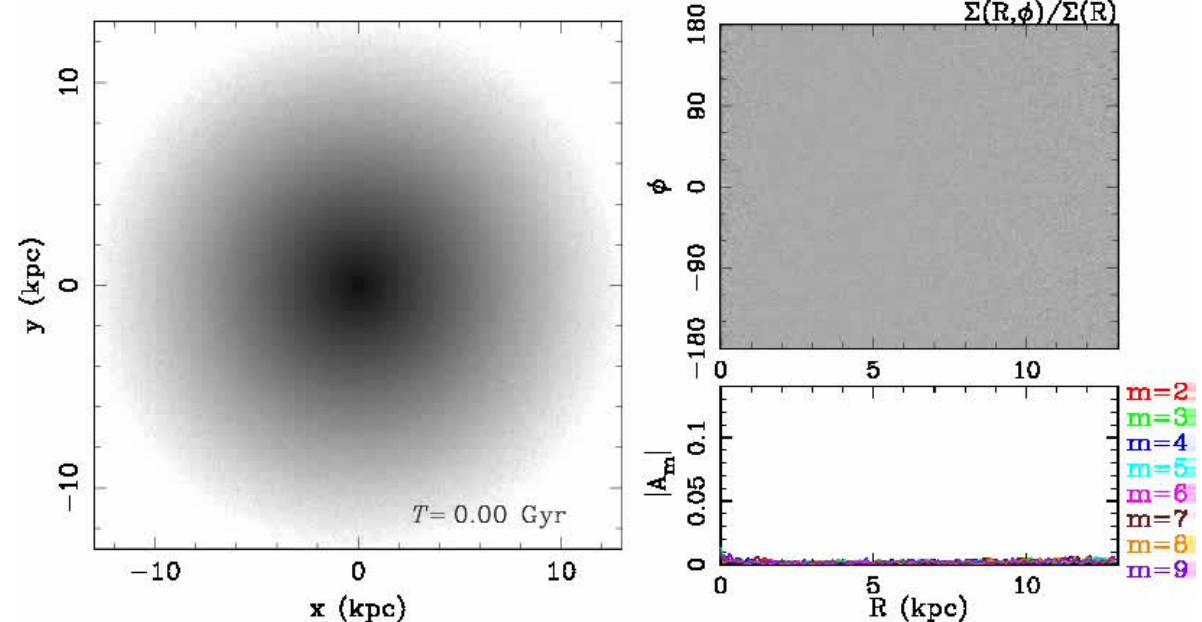
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Density wave theory
 (Lin & Shu 1964)



©Ingo Berg

Recurrent transient spiral
 (e.g. Goldreich & Lynden-Bell 1965)



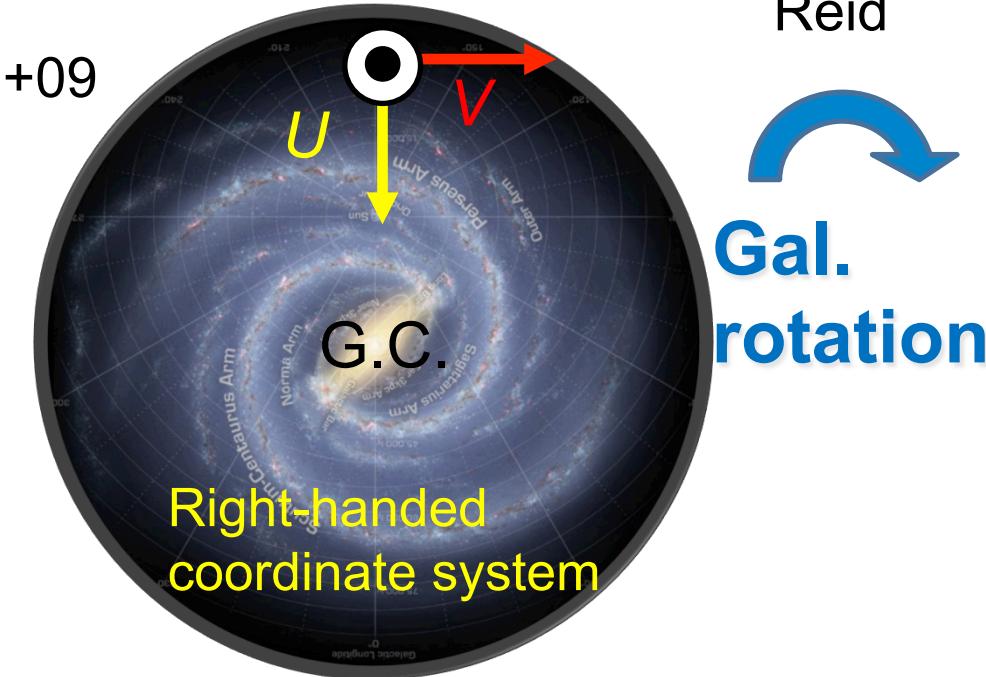
Fujii et al. 2011:N-body simulation

- Spiral arm is the most prominent structure in a disk galaxy.
- However, the nature, origin and evolution are still unknown.

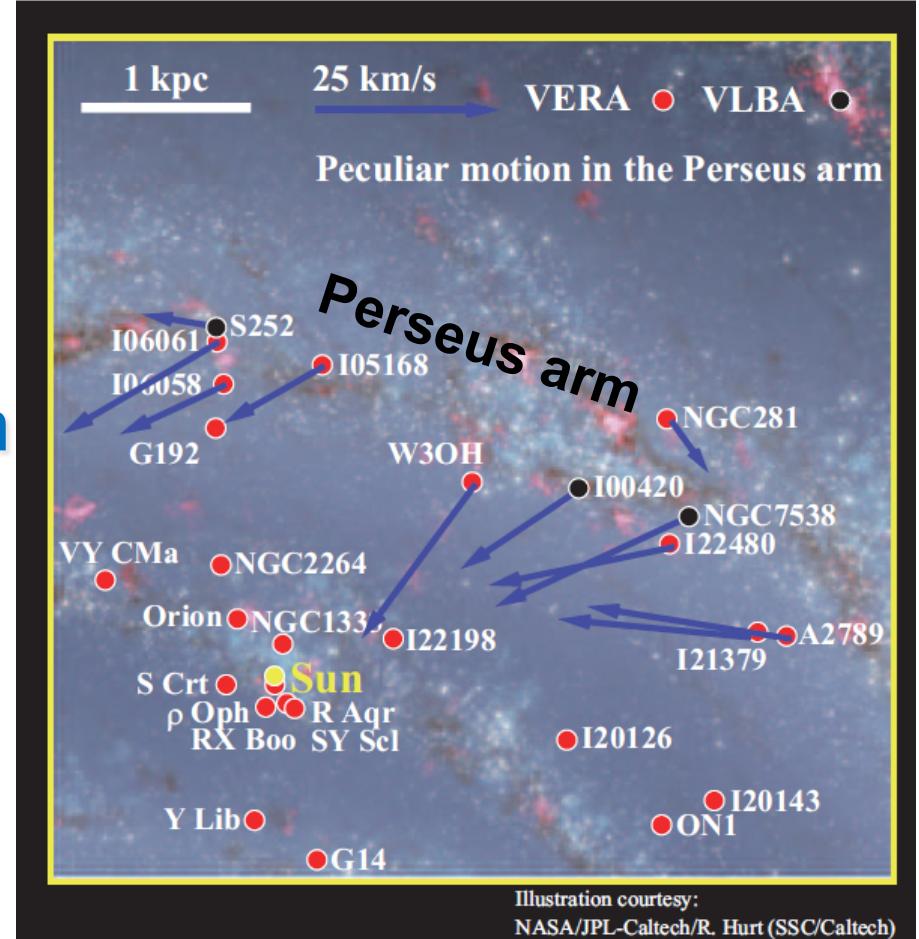
3-D astrometry results

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3D non-circular motion (U , V , W) in Reid



Choi et al. (2014) found the same tendency with 25 sources!



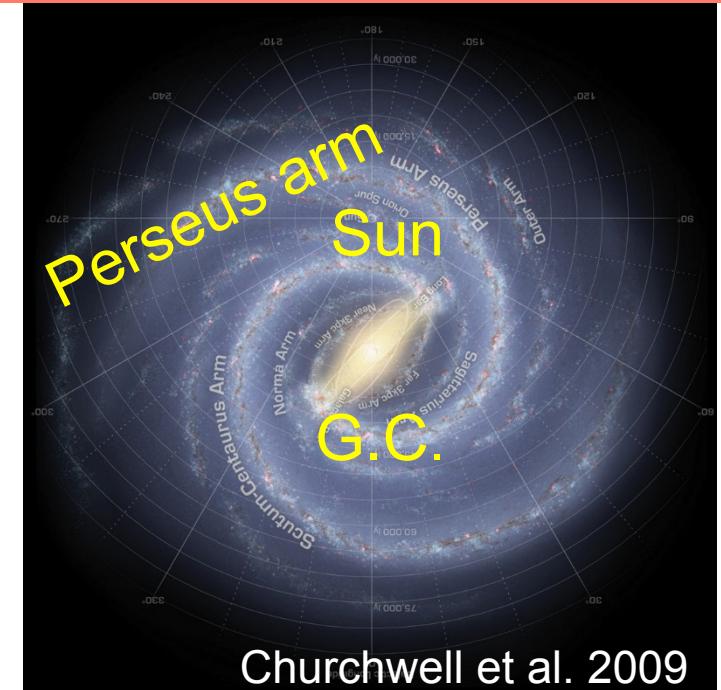
Averaged non-circular motions with 8 sources in the Perseus arm
 $(U_{\text{mean}}, V_{\text{mean}}) = (9.8 \pm 2.6, -17.8 \pm 2.5) \text{ km/s}$ in Sakai et al. (2012)
→ Systematic inward motion and slower Gal. rotation

Goals of this research

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① We compare astrometry (3D) results for the Perseus arm with an analytic gas dynamics model based on the density-wave theory.

② Using the results, we make a suggestion toward an incoming astrometry (e.g. Gaia).



Ultimate Goal

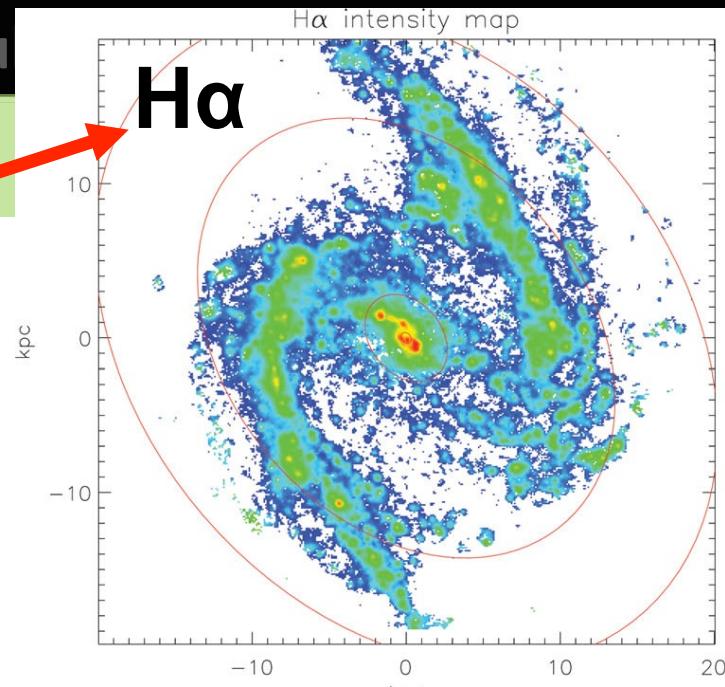
- We aim to understand the origin and evolution of the spiral arm in the Milky Way.

The least squares fit

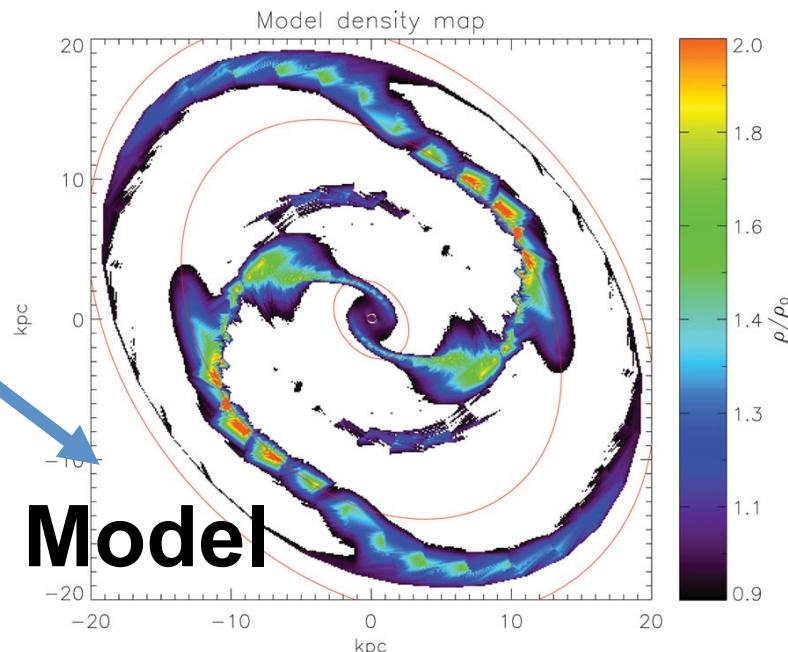
Observables
 (U, V) for 27 sources in the Perseus arm

χ^2_v was minimized

- Spiral model (density wave)**
- Pinol-Ferrer+12 & 14 (analytic)
 - Equations of motion with a gas friction term were solved using the linear approximation.
 - 7 model parameters



(Zanmer Sanchez+08)



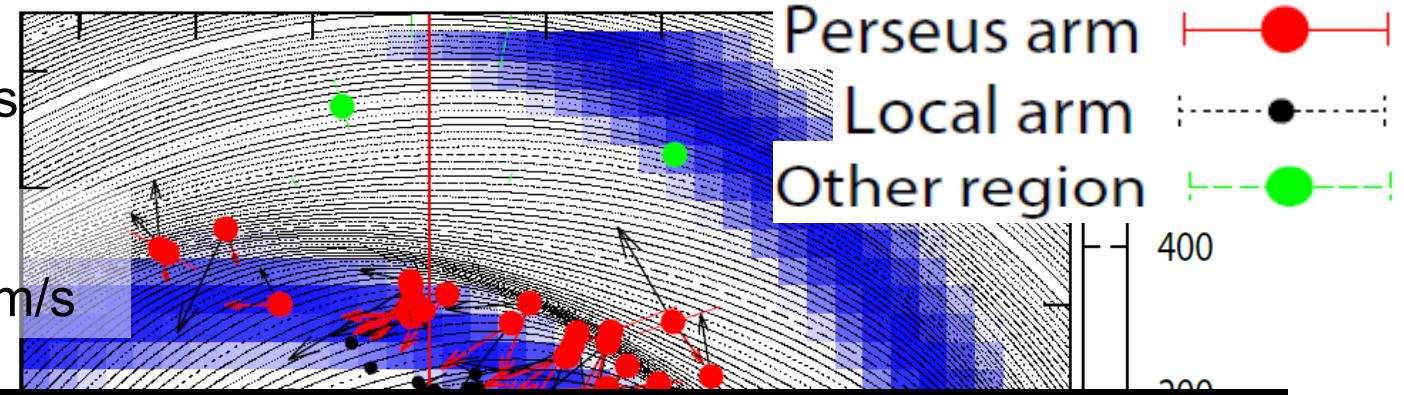
Model

Results

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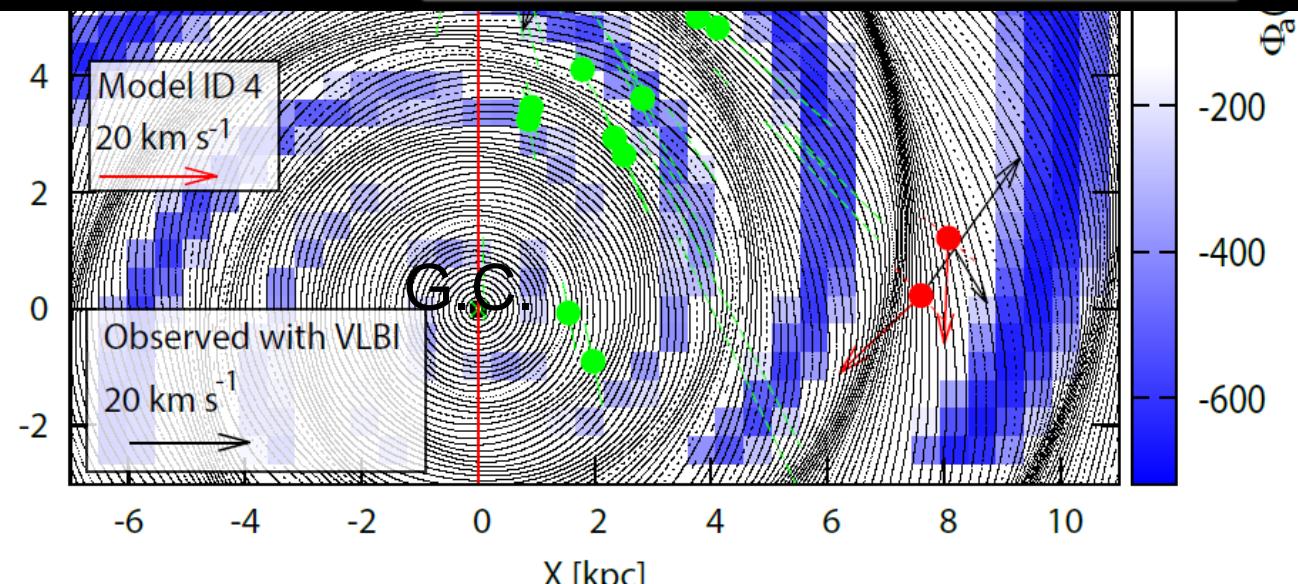
$$(U_{\text{mean}}, V_{\text{mean}}) = (8+/-3, -9+/-2) \text{ km/s}$$

$$(\Delta U_{\text{mean}}, \Delta V_{\text{mean}}) = (-2+/-2, -2+/-2) \text{ km/s}$$



Gas density mode An offset between spiral potential and gas might be checked by using **Gaia and VLBI results!**

Spiral potential model

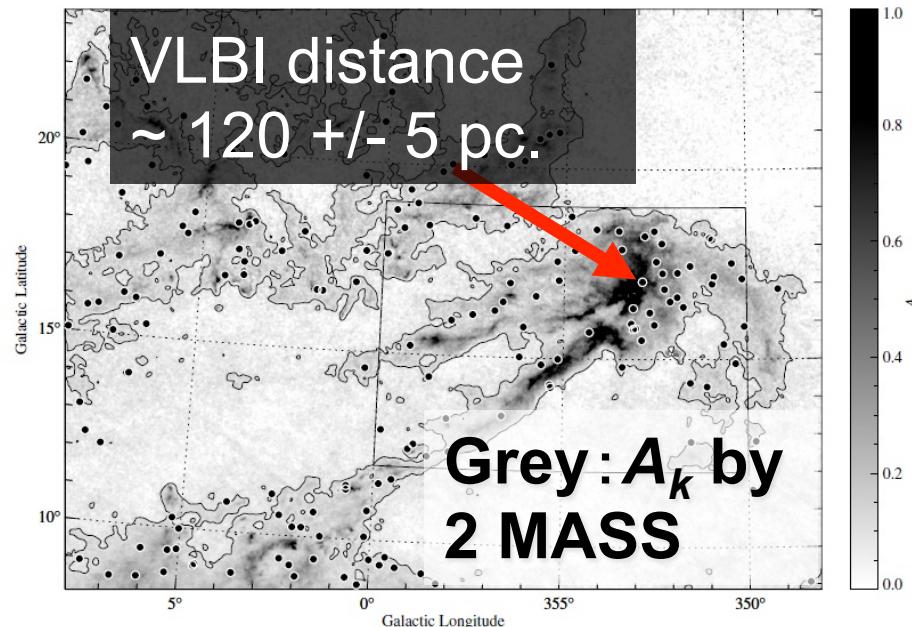


Collaboration between Gaia and VLBI

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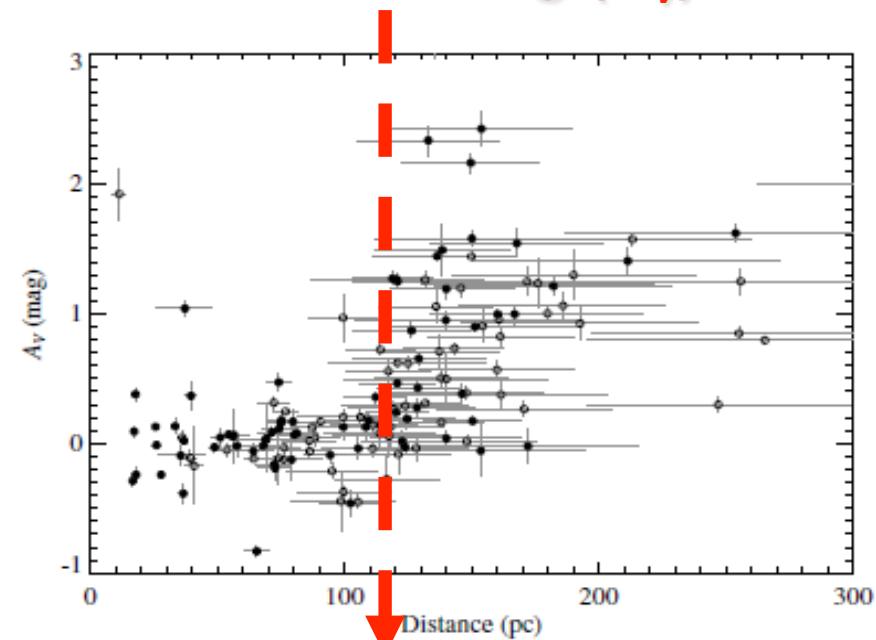
- ① Reference (ICRF & GCRF) frame ties
- ② Cross check (e.g. “Pleiades distance controversy”, Melis+14)
- ③ Model selection (e.g. Spiral arm)

Ophiuchus Region



● : *Hipparcos results*
 (Lombardi+08)

Reddening (A_v)



D ~ 120 pc

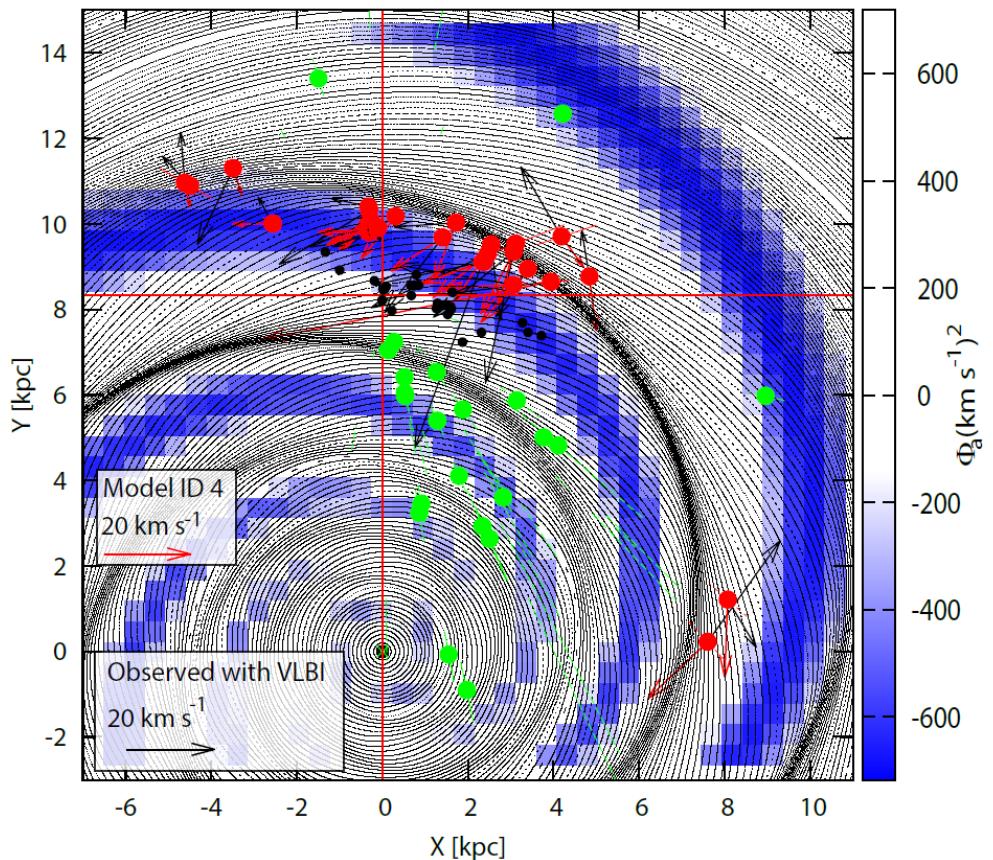
Conclusion

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- We tested the density-wave theory using VLBI astrometry results. Then, we found

①A dense gas region at the downstream of the spiral potential (model)

②Collaboration between Gaia and VLBI astrometry is crucial for model selection.



Perseus arm

Other regions

Local arm

VLBI astrometry vs. The model of the Milky Way